The Free Klein Gordon Field Theory

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Abstract

A single-particle relativistic theory turns out to be inadequate for many situations. Thus, we begin to develop a multi-particle relativistic description of quantum mechanics starting from classical analogies. We start with a Lagrangian description, and use it to build a Hamiltonian description we can then quantize. We then illustrate the decomposition of the eld into its positive and negative energy components, and de ne the creation and annihilation operators.

1 Introduction

Relativistic quantum mechanics turns out to be inadequately described by a single-particle theory. One of the primary reasons for this is that a relativistic particle may have enough energy to create other particles, and thus any theory describing it must account for this.

A theory which describes an in nite-degree-of-freedom system could be built out of a function de ned over allersystebuitoile5205tow8(era7(y)-4408(bsuc7(y)hTJ 0 -11.955 Td [(pais.ng-44 We will assume there exists an action S de ned as normal, which we can write in terms of L: 7 t 7 t 7 t

$$S = \int_{t_i}^{L} dt \, L = \int_{t_i}^{L} dt \, d^3 x \, L = \int_{t_i}^{L} d^4 x \, L$$
(2)

We will further assume that $\ensuremath{{\msc L}}$ depends only on the $% \ensuremath{{\msc eld}}$ eld variable and its $% \ensuremath{{\msc rst}}$ rst derivative, i.e.

$$L = L((x); @ (x))$$
(3)

Now we will change the eld arbitrarily and in nitesimally to nd what conditions will result in a stationary action

and apply the vanishing boundary condition

$$(\mathbf{x};t_i) = (\mathbf{x};t_f) = 0 \tag{5}$$

Now we can look at an in nitesimal change in S:

We can expand this using the total derivative of *L*:

$$= \frac{\sum_{t_f} t_f}{t_i} \mathcal{Q}^4 X \quad \frac{@ \bot}{@ (X)} \quad (X) + \frac{@ \bot}{@ @}$$

We will pick the Lagrangian density

$$L = \frac{1}{2} @ @ \frac{1}{2} m^{2} ^{2}$$
 (11)

This choice can be motivated by looking at the Lagrangian for a classical harmonic oscillator,

$$L = \frac{1}{2}mv^2 \quad \frac{1}{2}m!^2x^2 \tag{12}$$

We see that if we treat (x) as x, @ is analogous to v, and doing so gives us (11).

Evaluating both sides to be

$$\frac{@L}{@(x)} = m^2(x)$$
(13a)

$$\frac{@L}{@@} \frac{}{(x)} = \frac{@}{@@} \frac{1}{2} @ \qquad @ \tag{13b}$$

$$= \frac{1}{2} \qquad \frac{@}{@@} @ \qquad @ \qquad + @ \qquad \frac{@}{@@} @ \qquad (13c)$$

$$=\frac{1}{2} \qquad (@ + @) \qquad (13d)$$

$$=\frac{1}{2} \qquad (2@) \qquad (13e)$$

with the Hamiltonian itself de ned as

$$H = \int_{-\infty}^{\infty} d^3 x H \tag{17}$$

Clearly, these are de ned very similarly to their classical de nitions.

equations with :

$$-(x) = f(x); Hg$$

$$= (x); d^{\beta}y \frac{1}{2} (y) + \frac{1}{2}(r(y)) (r(y)) + \frac{1}{2}m^{2}$$

$$= \frac{1}{2}^{Z} d^{\beta}x f(x); (y)g_{x^{0}=y^{0}}$$

$$= d^{\beta}y (y)f(x); (y)g_{x^{0}=y^{0}}$$

$$= d^{\beta}y (y; x^{0})f(x); (y)g_{x^{0}=y^{0}}$$

$$= d^{\beta}y (y; x^{0})f(x); (y)g_{x^{0}=y^{0}}$$

$$= (x)$$
(25)

Here we have used that the Hamiltonian is time-independent, so we can assert that $x^0 = y^0$, allowing us to use the equal-time relation (19). We can go through a similar process to nd -:

$$\begin{aligned} -(x) &= f(x); Hg \\ &= (x); d^{8}y \frac{1}{2} {}^{2}(y) + \frac{1}{2}(r(y)) (r(y)) + \frac{1}{2}m^{2} {}^{2} \\ &= d^{8}y(r(y) f(x); r(y)g_{x^{0}=y^{0}} \\ &+ m^{2}(y)f(x); {}^{2}(y)g_{x^{0}=y^{0}}) \\ &= d^{8}y(r(y;x^{0}) r({}^{3}(x y)) \\ &+ m^{2}(y;x^{0})({}^{3}(x y)) \\ &+ m^{2}(x) = r^{2}(x) m^{2}(x) \end{aligned}$$
(26)

These rst-order Hamilton equations give us the second-order equation

•
$$(x) = -(x) = r^2 (x) m^2 (x)$$

! $(\cdot r^2 (x)) + r^2 X$ equations

4 Field Decomposition

The following plane wave equation set forms a complete basis for solutions to the Klein-Gordon equation [1]:

$$(X) = e^{-ik \cdot X} \tag{30}$$

We can use this basis to expand in this basis: 7

$$(x) = C d^{4}k e^{-ik \cdot x} (k) ; \quad C = \frac{1}{(2)^{\frac{3}{2}}}$$
(31)

This is essentially a Fourier transform of $\tilde{c}(k)$, with *C* introduced for later convenience. We can now plug this into the Klein-Gordon equation:

$$C(@ @ + m^{2}) \overset{L}{d^{4}k} e^{-ik \times -}(k) = 0$$

$$Z$$

$$! C d^{4}k (@ @ + m^{2})e^{-ik \times -}(k) = 0$$

$$Z$$

$$! C d^{4}k (k^{2} + m^{2})^{-}(k)e^{-ik \times -} = 0$$
(32)

So at least one of the following must be true:

$$k^2 = m^2 \tag{33}$$

$$\tilde{k} = 0 \tag{34}$$

So, ~ is only non-vanishing when $k^2 = m^2$; we'll de ne ~ in the following way to capture this:

$$^{\sim}(k) = {}^{SO_{i}}$$

Now we can plug this new form into (35b): Z

$$(x) = C dk^{0} d^{3}k \frac{1}{2E_{k}} (k^{0} E_{k}) + (k^{0} + E_{k}))$$

$$= C d^{3}k \frac{1}{2E_{k}} e^{i(E_{k}x^{0} + \mathbf{k}\cdot\mathbf{x})} a(E_{k};\mathbf{k}) + e^{i(E_{k}x^{0} + \mathbf{k}\cdot\mathbf{x})} a(E_{k};\mathbf{k}) + e^{i(E_{k}x^{0} + \mathbf{k}\cdot\mathbf{x})} a(E_{k};\mathbf{k})$$
(39)

We will replace E_k with k^0 and switch $\mathbf{k} \mathbf{Z}_0 = \mathbf{k}$ in the second term to obtain

$$(x) = C \int_{-\infty}^{\infty} d^3k \frac{1}{2E_k} x^{3}$$

References

- [1] Ashok Das. Lectures on Quantum Field Theory. World Scienti c, 2008.
- [2] Ben Saltzman. Second quantization of the klein-gordon equation. 2018.